

Large Area Continuous Position Sensitive Diamond Detector: First tests

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Outline

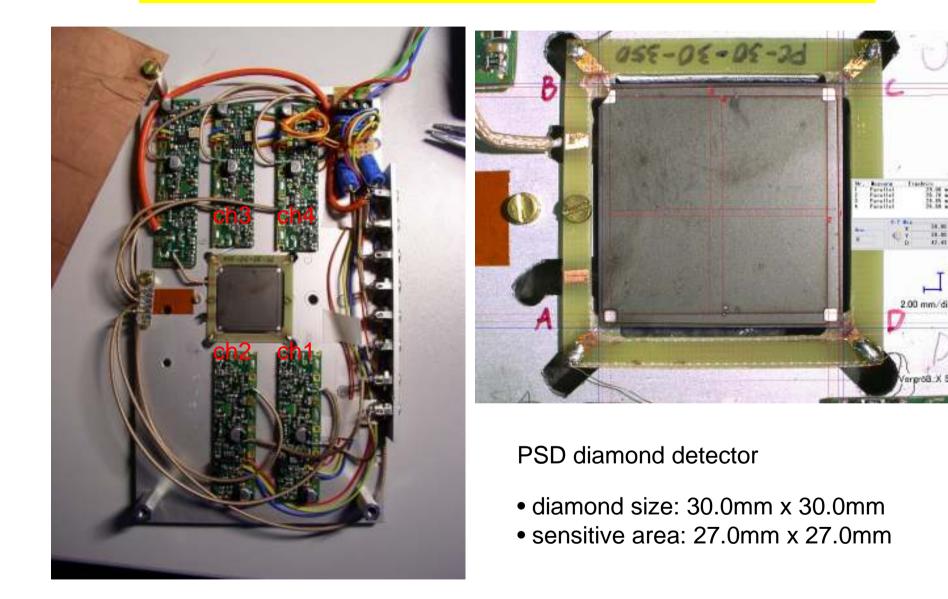


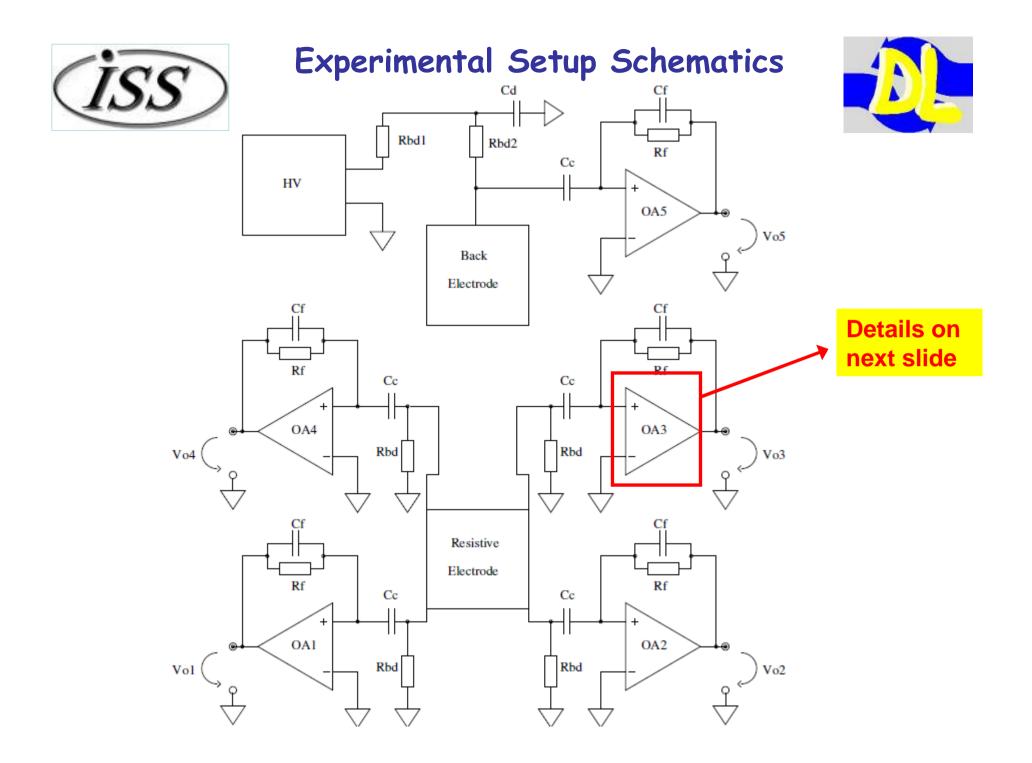
Preparatory work for understanding how to make a Large Area Position Sensitive Diamond Detector.

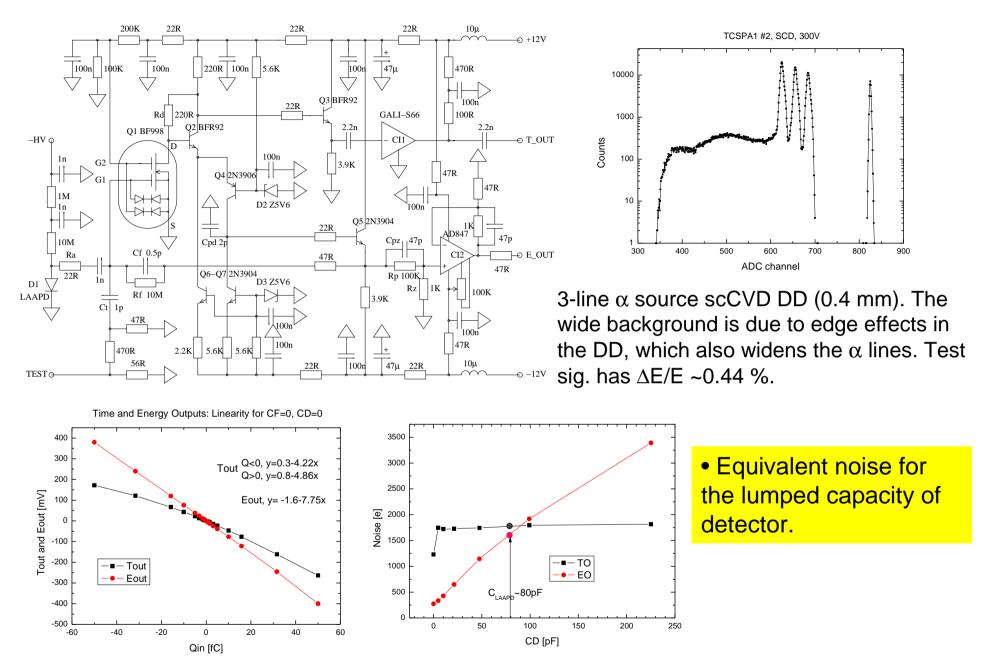
The final target is to use the new Diamond on Iridium material.

- 1. Set-up description
- 2. In-beam tests
- 3. Summary and outlook

Position Sensitive Detector – Overview

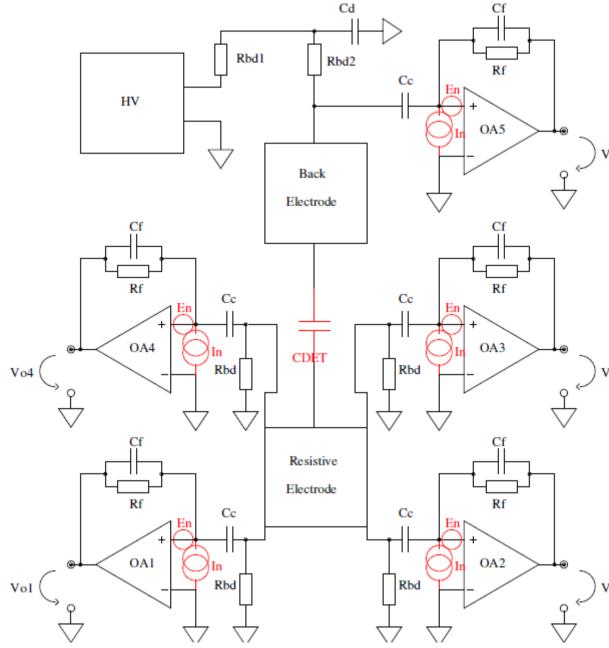






M. Ciobanu et all., A Charge Sensitive Amplifier for Time and Energy Measurements IEEE-Nuclear Science Symposium, Conf. Rec., pp. 2028-2032, 2008

Schematics Including the CSA Noise Generators



• The ideal CSA has 0Ω input impedance. The real one has $\sim k\Omega$.

^{vo5} • The resistance of the resistive electrode and the capacitance C_{DET} connects together all amplifiers. The noise generators En and In contribute over the whole network and the first
 vo3 connection test showed a factor of 7 increase in the

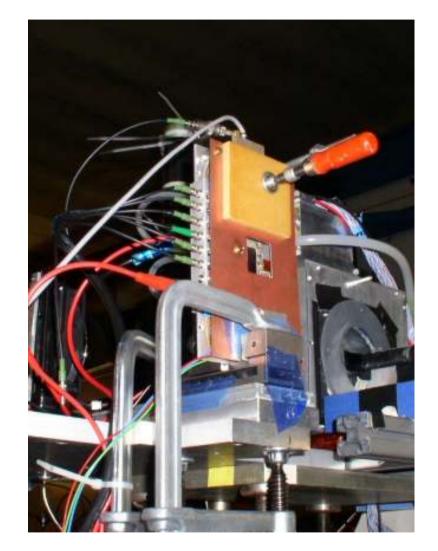
- factor of 7 increase in the noise of the CSA output.
- In order to decrease this effect, we have reduced Rf and increased Cf, with the drawback of additional loss in responsivity and resolution.

Ni-58 @ 1.7GeV/u??

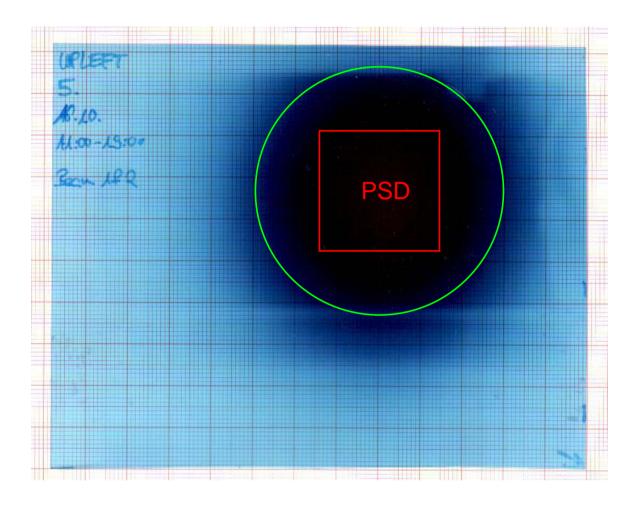
Beamtimes in CaveB (FOPI)

16.10. - 19.10.12 (high rate)

Trigger : Back electrode

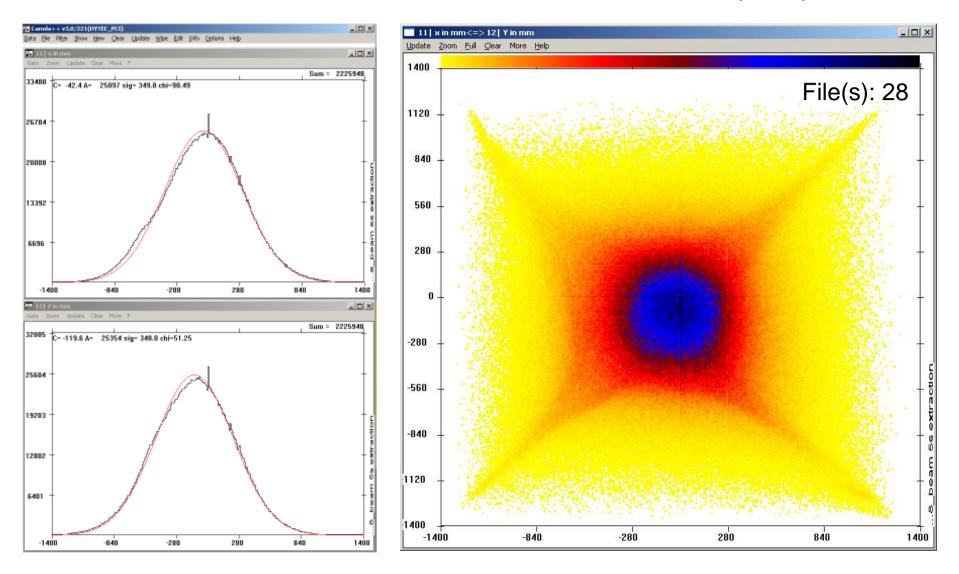


Beam



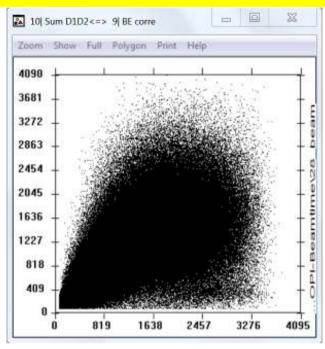
- beam: Ni-58 @ 1.7GeV/u
- detector position: beam center, ~2m before beam dump in FOPI Cave (B)
- beam diameter at detector position ~60mm

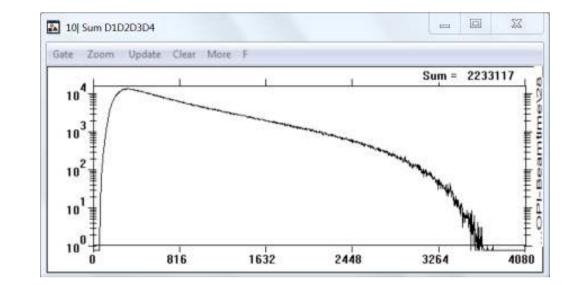
Gate from PSD Back Electrode (BE)



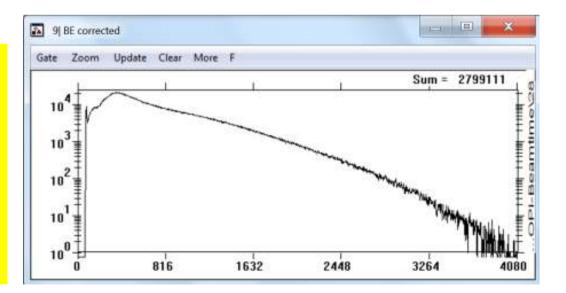
- not pin-cushion corrected spectra of x- , y-position
- no amplitude filter

Events depending on energy – all energies

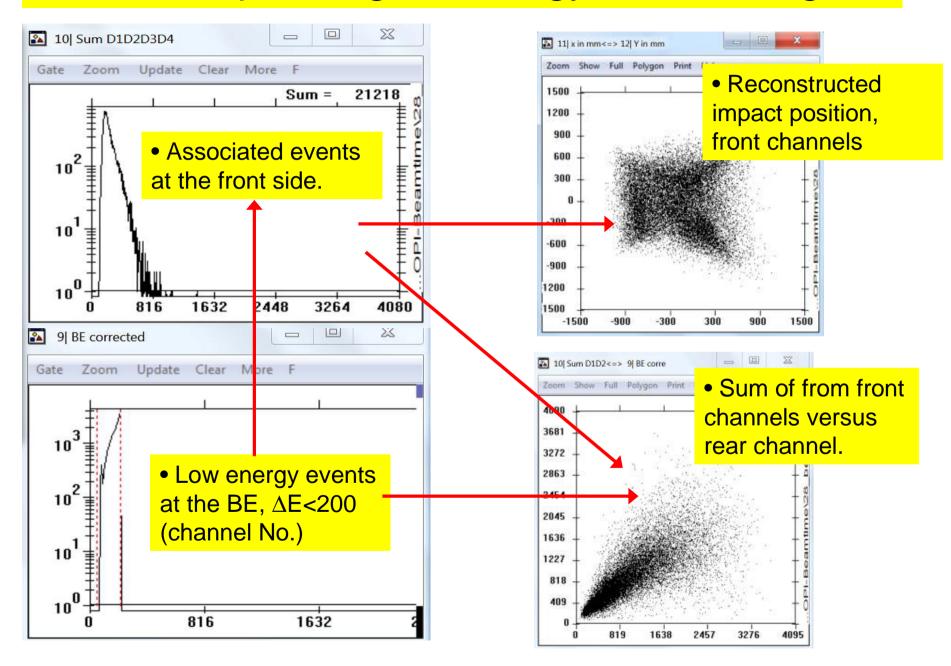


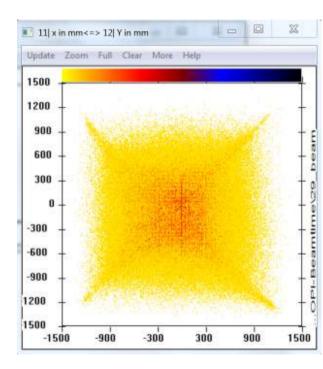


- Right: Histogram of events depending on energy deposition (top=sum of front channels, bottom=rear chanel)
- Left: Scatter plot of rear channel versus sum of the front channel
 ideally, the points should line up along the diagonal

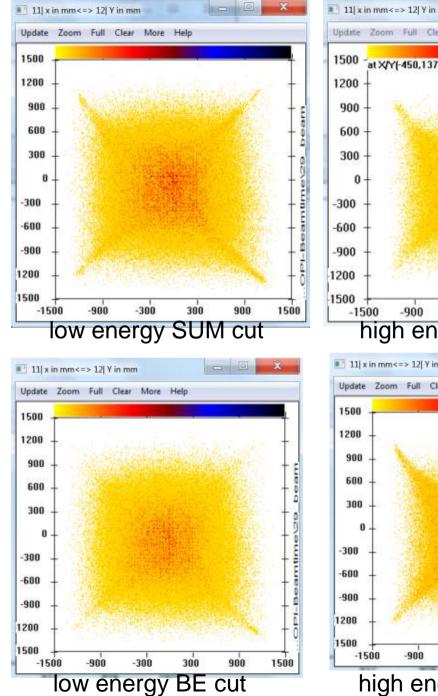


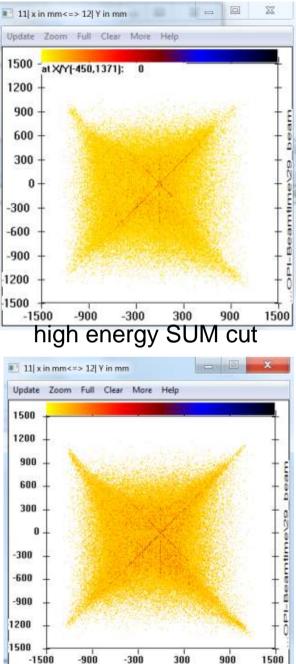
Events depending on energy – low energies





All 29_data, no cut





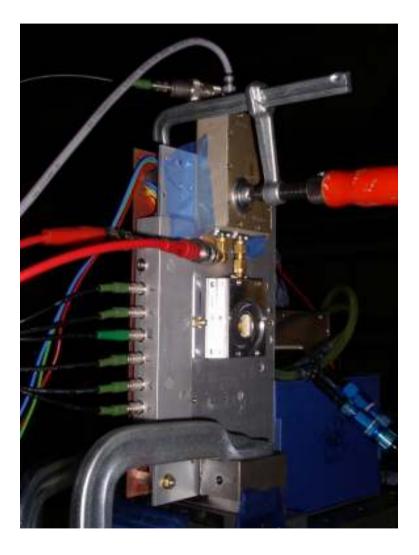
high energy BE cut

Ni-58 @ 1.7GeV/u??

Beamtimes in CaveB (FOPI)

29.10. - 5.11.12 (low rate)

Trigger: pcCVD 8mm or scCVD 1.2mm

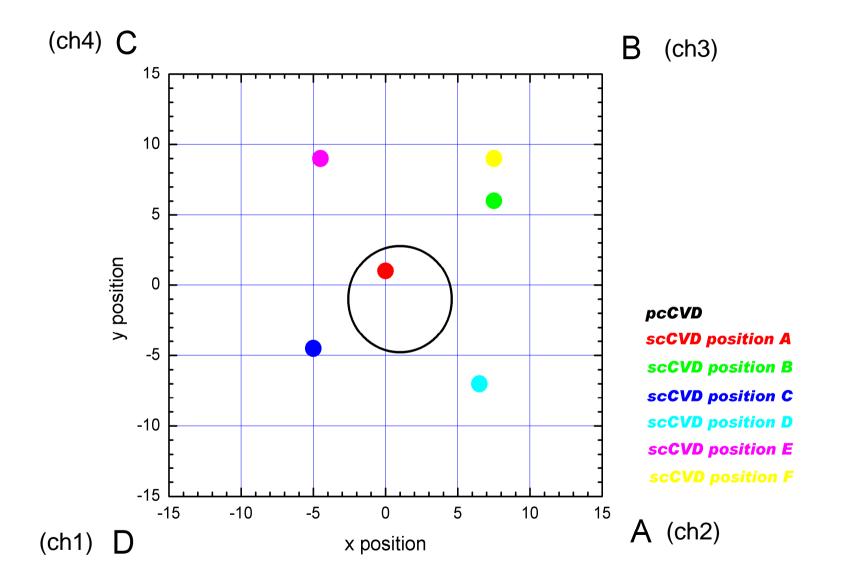


scCVD diamond 10B50 Gate detector

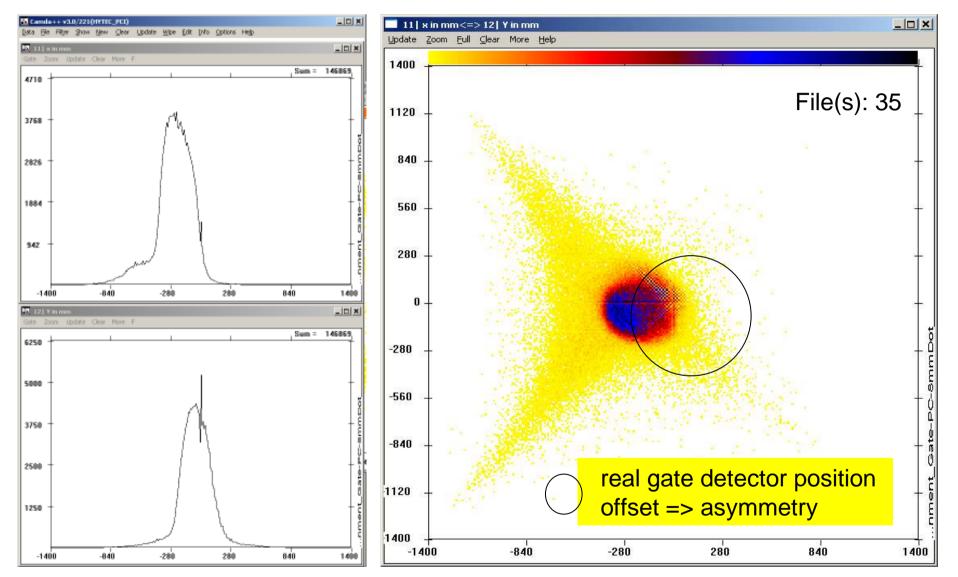


• 1.25mm circular electrode

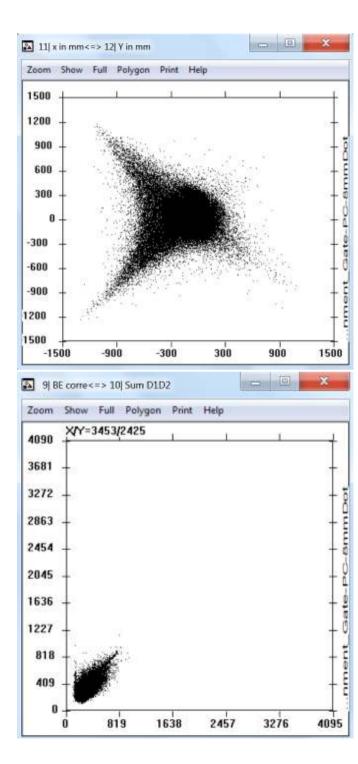
Gate detector positions relative to PSD

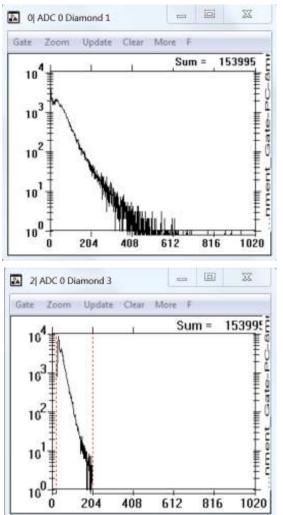


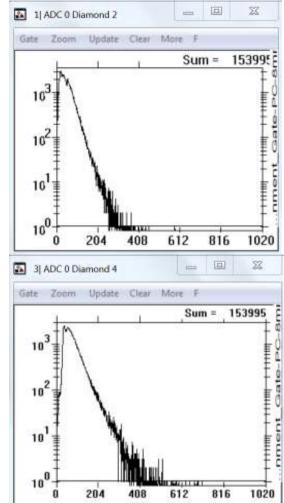
Gate from pcCVD 210um with circular 8mm electrode

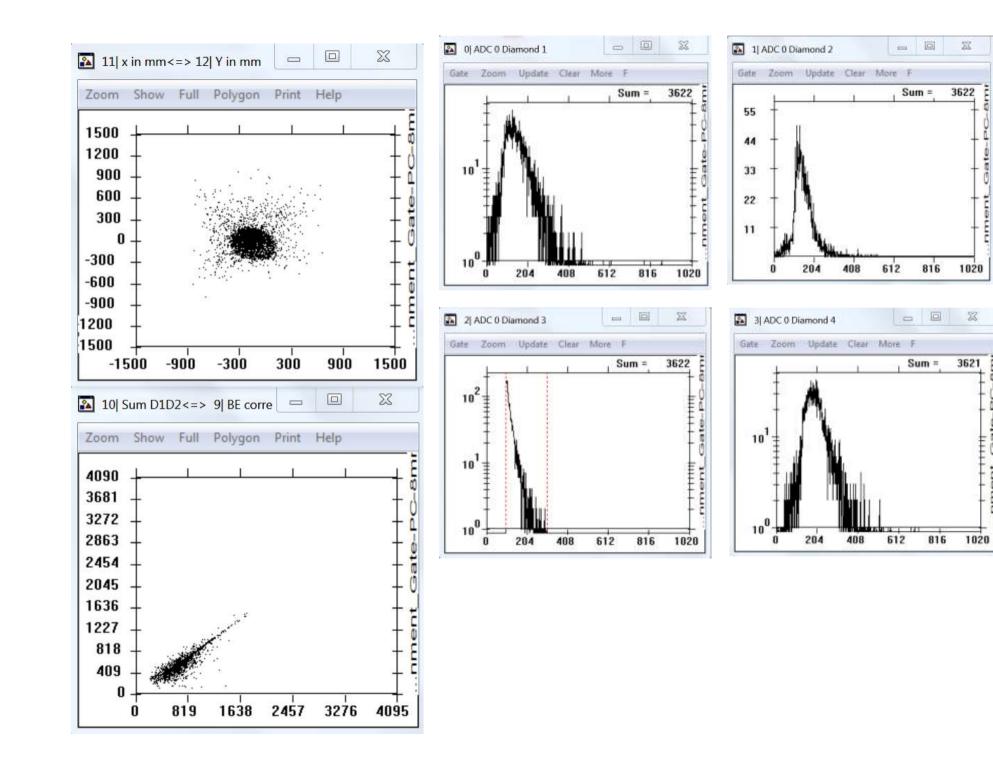


- not pin-cushion corrected spectra of x- , y-position
- no amplitude filter



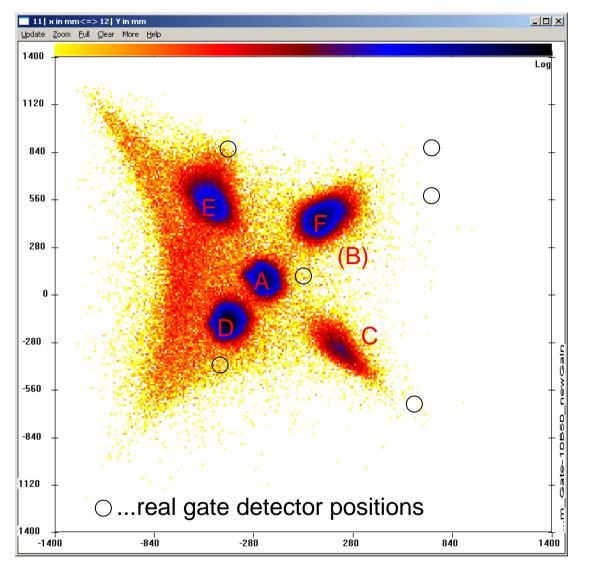






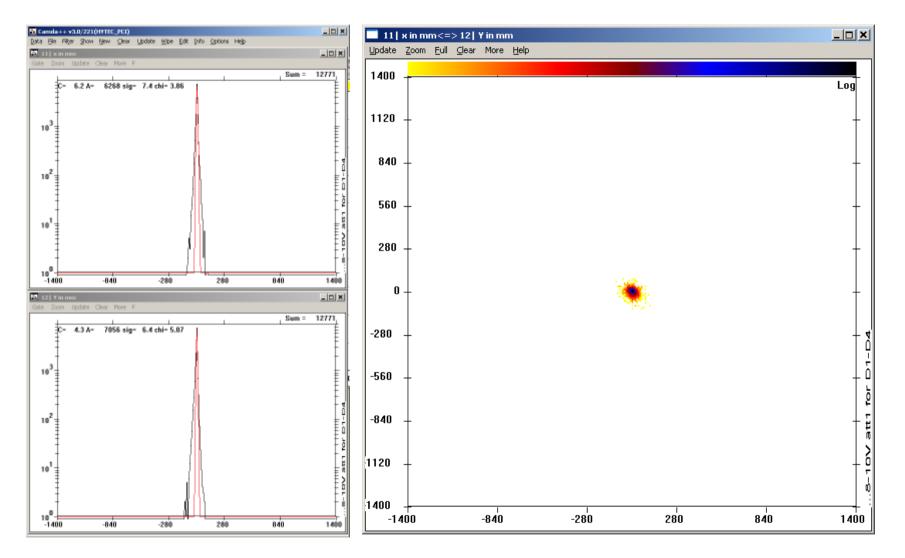
nment

Gate from scCVD 10B50 with circular 1.25mm electrode for different relative positions (A-F)



- not pin-cushion corrected spectra of x-, y-position
- amplitude filter: sum of all 4 channels within +/-5% of BE signal

Pulser calibration



- no filter
- all pulser signals (same amplitude for all channels) after correction at "center position")
- Message: the FEE cannot be responsible for the effects observed





Summary

- The first LACPSDD experimental set-up has been made and tested.
- The main results obtained in the described tests confirms that the detector has a highrate capability. Due to the distributed resistance and capacitance, the lumped equivalent scheme of the detector should use resistor and capacitor with values less then R□ and CDET.
- The 5 CSA used have an significant increase of the noise level when connected to the detector. We have decreased the feedback resistance and increased the feedback capacitance in order to decrease the crosstalk between channels, with additional decrease of the charge responsivity and increase in the noise level.
- We have tried to use low noise broadband amplifiers with the same set-up. The cabling of the experimental set-up was not sufficiently clean to avoid oscillations.
- The in-beam tests have shown two types of signals:
 1) lower amplitudes signals which give a position error with cushion shape
 2) higher amplitudes signals which give a random position error
 We do not know what is the origin of the two types of signal





Outlook 1

- For the next step, the most important task will be to change the amplifier type to avoid the crosstalk between channels and to have enough charge responsivity to measure alpha particles. The picking time <10 ns.
- We must validate the 3D simulation environment needed for the optimization of the experimental set-up, in particular to take into account the distributed character of the electrical parameters.
- For in-beam tests we need a position etalon.
 - With a 10mm x10mm pcCVD electronic grade from the last productions of Element 6, one could make a 2D duo-lateral LACPSDD with very small position error.
- The comparative tests with alpha particles of the two types of LAPSDD can give the optimum algorithm for the minimal systematic error of position reconstruction. After inbeam tests we shall choose the final structure.
- At that moment we shall be ready to finally use the DoI material.

Outlook 2 : Prospects for Application to Space Experiments

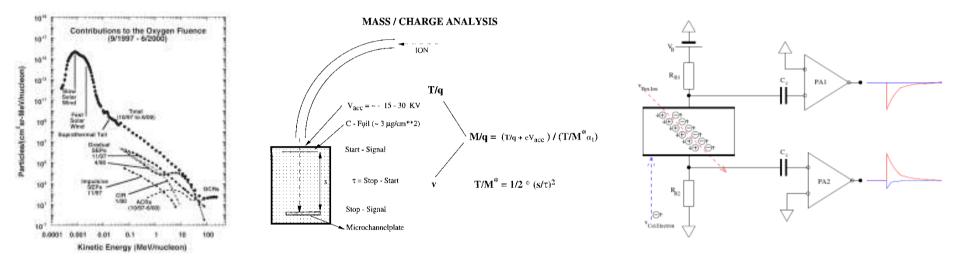


Figure 1. Fluence of energetic oxygen nuclei from solar wind to galactic cosmic ray energies measured by the Advanced Composition Explorer (ACE) satellite. From *Mewaldt et al. (2006)*.

Figure 2. Schematic view of a mass spectrometer, including a curved plate electrostatic analyzer and a time-of-flight section (From *Klecker, 2001*).

Figure 3. Electrical scheme that allows the rejection of penetrating radiation, by collecting the signal from both sides of the DD in conjunction with microchannel plates, in particle spectrometry.







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